

Branching diffusion representation for nonlinear Cauchy problems and Monte Carlo approximation

Pierre Henry-Labordère*

Nizar Touzi†

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Abstract

We provide a probabilistic representations of the solution of some semilinear hyperbolic and high-order PDEs based on branching diffusions. These representations pave the way for a Monte-Carlo approximation of the solution, thus bypassing the curse of dimensionality. We illustrate the numerical implications in the context of some popular PDEs in physics such as nonlinear Klein-Gordon equation, a simplified scalar version of the Yang-Mills equation, a fourth-order nonlinear beam equation and the Gross-Pitaevskii PDE as an example of nonlinear Schrödinger equations.

1 Introduction

Similar to the intimate connection between the heat equation and the Brownian motion, linear (second-order) parabolic partial differential equations are connected to stochastic processes. More precisely, the Feynman-Kac formula states that a linear parabolic PDE with infinitesimal generator $\mathcal{L} := b \cdot \partial_x + \frac{1}{2} \text{Tr}[\sigma \sigma^T \partial_{xx}^2]$, and terminal condition $f_1(\cdot)$ at time T , can be written as a conditional expectation of $f_1(X_T)$ involving the Itô process X associated with the generator \mathcal{L} . This connection allows to devise numerically approximations of the solution of such PDEs by probabilistic (Monte Carlo) methods, which represents a clear advantage in high dimensional problems as the error estimate induced by the central limit theorem is dimension-free.

An important focus was put on the extension to nonlinear (second-order) parabolic PDEs, see e.g. [12] for a quick review of existing methods.

- A first attempt was achieved by exploiting the stochastic representation by means of backward stochastic differential equations, see Bally & Pagès [3], Bouchard & Touzi [6],

*Société Générale, Global Market Quantitative Research, pierre.henry-labordere@sgcib.com

†Ecole Polytechnique Paris, Centre de Mathématiques Appliquées, nizar.touzi@polytechnique.edu. This work benefits from the financial support of the ERC Advanced Grant 321111, and the Chairs *Financial Risk* and *Finance and Sustainable Development*.

Zhang [20], and the extension to the fully nonlinear setting by Fahim, Touzi & Warin [11]. The induced numerical method involves repeated computations of conditional expectations, resulting in a serious dimension dependence of the corresponding numerical methods. In fact, as highlighted in [11], this method can be viewed as part of the traditional finite-elements algorithm.

- More recently, Henry-Labordère [13] suggested a numerical method based of an extension of the McKean [19] branching diffusion representation of the so-called KPP equation (Kolmogorov-Petrovskii-Piskunov) to a class of semilinear second order parabolic PDEs with power nonlinearity in the value function u . The resulting algorithm is purely probabilistic. In particular its convergence is controlled by the central limit theorem, and is independent of the dimension of the underlying state. The validity of this method in the path-dependent case, and for further analytic nonlinearities which are of the power type in the triple $(u, \partial_x u, \partial_{xx}^2 u)$ is analyzed in [13, 14, 15, 16]. We also refer to Agarwal and Claisse [1] for the extension to elliptic semilinear PDEs, and Bouchard, Tan, Warin & Zou [5] for Lipschitz nonlinearity in the pair $(v, \partial_x v)$. A critical ingredient for the extension is the use of Galton-Watson trees weighted by some Malliavin automatic differentiation weights.

Our objective in this paper is to show that the above branching diffusion approach extends to more general Cauchy problems, including hyperbolic and higher order PDEs. Such an extension may seem to be unexpected due to the intimate connection between parabolic second order PDEs and diffusions generators. However, probabilistic representations of some specific examples of hyperbolic PDEs did appear in the previous literature. The first such relevant work traces back to Kac [18] in the context of the one-dimensional telegrapher equation:

$$\partial_{tt}u - c \partial_{xx}u + (2\beta)\partial_t u = 0, \quad u(0, x) \equiv f_1(x), \quad \partial_t u(0, x) \equiv f_2(x),$$

where β and c are constant parameters, and the boundary data f_1, f_2 are some given functions. Observing that we may express $u(t, x) = \frac{1}{2}[u_0 + u_1](t, x)$, where the pair (u_0, u_1) solves the coupled system of first order PDEs:

$$\partial_t u_i + (-1)^i c \partial_x u_i - \beta(u_{1-i} - u_i) = 0, \quad u_i(0, x) = f_1(x) - \frac{(-1)^i}{c} \int_0^x f_2(y) dy, \quad i = 0, 1,$$

A branching mechanism representation of u_0 and u_1 is obtained by following Mc Kean's representation for (interacting) KPP equations (with zero diffusion).

We next mention the work of Dalang, Mueller & Tribe [8] who introduced an alternative stochastic representation for a class of *linear* Cauchy problems $Lu = Fu$, with a potential F , including the linear wave equation. Their starting point is the well-known representation

$$u(t, x) = w(t, x) + \int_0^t \int_{\mathbb{R}^d} V(t-r, x-y) u(t-r, x-y) S(r, dy) dr, \quad (1.1)$$

where w is the solution with zero potential, and S is a fundamental solution, restricted to be representable by a signed measure with $\sup_t |S(t, \mathbb{R}^d)| < \infty$. By substituting formally

$u(t-r, x-y)$ on the right hand side by the last expression of u , one formally obtains a candidate representation for u as $\sum_{m \geq 0} H_m(t, x)$, where $H_0 = w$, and $H_m = \int_0^t \int_{\mathbb{R}^d} V(t-r, x-y) H_{m-1}(t-r, x-y) S(r, dy) dr$, $m \geq 1$. Finally, by convenient normalization of the kernel S , the last expression induces a probabilistic representation.

Subsequently, Bakhtin and Mueller [2] considered the one-dimensional nonlinear wave equation

$$\partial_t^2 u - \partial_x^2 u = \sum_{j \geq 0} a_j u^j \quad \text{on } \mathbb{R}_+ \times \mathbb{R}, \quad u(0, \cdot) = f_1, \quad \partial_t u(0, \cdot) = f_2 \quad \text{on } \mathbb{R},$$

and obtained a probabilistic representation by means of stochastic cascades, which mimics exactly the McKean branching diffusion representation of the KPP equation [19], similar to [13, 14]. Our main contribution in this paper is to show that such a representation holds for a wider class of Cauchy problems, in arbitrary dimension, and with analytic nonlinearity in $(u, \partial_x u)$.

Our starting point is that the representations of [8] and [2] are closely related to the McKean [19] representation of KPP equations, and the corresponding extensions in [13, 14, 16]. This in fact opens the door to a much wider extension reported in Section 2 in the context of linear Cauchy problems with constant coefficients, and in Section 3 in the context of nonlinear Cauchy problems with constant coefficients, and analytic nonlinearity in the value function u . The crucial tool for our extension is the so-called Duhamel formula which expresses the solution of such an equation as the convolution (i.e. integration) of the boundary conditions with respect to a family of fundamental solutions. This is in contrast with (1.1) which uses the single fundamental solution S , and requires that the solution w of the zero-potential equation be given. Then, the probabilistic representation appears naturally after convenient normalization of the fundamental solutions.

It is also remarkable that the Malliavin automatic differentiation technique, exploited in [15, 16] in order to address semilinear parabolic second order PDEs, extends naturally to the context of general Cauchy problems by introducing the space gradient of the fundamental solution. This observation is the key-ingredient for our extension in Section 4 to a general class of semilinear Cauchy problems with analytic nonlinearity in $(u, \partial_x u)$.

The performance of the Monte Carlo numerical method induced by our representation is illustrated in Section 5 on some relevant examples from mathematical physics. We start with two examples of semilinear wave equations: the Klein-Gordon equation in dimensions 1,2,3, which has a power nonlinearity in u , and a simplified version of the Yang-Mills equation (in dimension 1), which contains a nonlinearity in the space gradient. We observe that due to some restricting conditions which will be detailed in Sections 2, 3, and 4, our Monte Carlo approximation method does not apply to the multi-dimensional wave equation with space gradient nonlinearity. We next report some numerical experiments in the context of the nonlinear one-dimensional Beam equation which contains two derivatives in time and four space derivatives. We finally consider the Gross-Pitaevskii equation in dimensions 1,2, and 3, as an example of nonlinear Schrödinger equation. All of the numerical results reveal an excellent performance of our Monte Carlo approximation method.

We finally emphasize that, throughout the paper, we consider the Cauchy problem on $\mathbb{R}_+ \times \mathbb{R}^d$, thus ignoring the important case of restricted domain $\mathcal{D} \subset \mathbb{R}^d$ for the space variable. We mention that Chatterjee [7] proved that the function $u(t, x) = \mathbb{E}_x[f(tX + \sqrt{\tau}Z, B_\tau)]$, with independent r.v. X, Z with standard Cauchy and normal distributions, respectively, and $\tau := \inf\{t > 0 : B_t \notin \mathcal{D}\}$ is the first exit time of an independent Brownian motion from the domain \mathcal{D} , solves the wave equation on $\mathbb{R}_+ \times \mathcal{D}$. However, this does not provide a representation for the wave equation on a restricted space domain as the determination of f from given $f_1(x) := u(0, x)$ and $f_2(x) := \partial_t u(0, x)$ is not transparent.

2 Probabilistic representation for linear Cauchy problems

2.1 Non-homogeneous Cauchy problem

For a smooth function $\phi : \mathbb{R}_+ \times \mathbb{R}^d \rightarrow \mathbb{R}$, we denote $\partial_t^0 \phi = D^0 \phi = \phi$, and

$$\partial_t^j \phi := \frac{\partial^j \phi}{\partial t^j} \quad \text{and} \quad D^\alpha \phi := \frac{\partial^{|\alpha|} \phi}{\partial x_1^{\alpha_1} \dots \partial x_d^{\alpha_d}}, \quad \text{for all } j \geq 1, \alpha \in \mathbb{N}^d.$$

Given two integers $N, M \geq 1$, we denote $\mathbb{N}_M^d := \{\alpha \in \mathbb{N}^d : |\alpha| \leq M\}$, and we consider some scalar parameters $(a_j)_{1 \leq j \leq N}, (b_\alpha)_{\alpha \in \mathbb{N}_M^d} \subset \mathbb{R}$ with

$$a_N = 1 \quad \text{and} \quad \{\alpha \in \mathbb{N}_M^d : |\alpha| = M \text{ and } b_\alpha \neq 0\} \neq \emptyset.$$

Throughout this paper, we consider nonlinear partial differential equations defined by means of the following linear Cauchy problem:

$$\sum_{n=1}^N a_n \partial_t^n u - \sum_{\alpha \in \mathbb{N}_M^d} b_\alpha D^\alpha u - F = 0 \quad \text{on } \mathbb{R}_+ \times \mathbb{R}^d, \quad (2.1)$$

$$\partial_t^{n-1} u(0, \cdot) = p_n f_n \quad \text{on } \mathbb{R}^d, \quad n = 1, \dots, N, \quad (2.2)$$

where the boundary data and the source term satisfy the following conditions:

$$f_n : \mathbb{R}^d \rightarrow \mathbb{R}, \quad 1 \leq n \leq N, \quad \text{and} \quad F : \mathbb{R}_+ \times \mathbb{R}^d \rightarrow \mathbb{R}, \quad \text{are bounded continuous,} \quad (2.3)$$

and p_1, \dots, p_N are scalar parameters in the simplex:

$$p_n > 0 \quad \text{for all } n = 1, \dots, N, \quad \text{and} \quad p_1 + \dots + p_N = 1.$$

2.2 Duhamel's formula for linear Cauchy problems

We next recall the general solution of non-homogeneous Cauchy problems with constant coefficients. We first introduce the \mathbb{C}^N -valued function with components $\hat{g} := (\hat{g}_1, \dots, \hat{g}_N)$:

$$\hat{g}(t, \xi) := (2\pi)^{-\frac{d}{2}} e^{tB(\xi)^T} \mathbf{e}_1, \quad t \geq 0, \quad \xi \in \mathbb{R}^d, \quad (2.4)$$

where (e_1, \dots, e_N) is the canonical basis of \mathbb{R}^N ,

$$B(\xi) := \left(\begin{array}{c|ccc} 0 & & & \\ \vdots & & I_{N-1} & \\ 0 & & & \\ \hline b(\xi) & -a_1 & \cdots & -a_{N-1} \end{array} \right), \text{ and } b(\xi) := \sum_{\alpha \in \mathbb{N}_M^d} i^{|\alpha|} b_\alpha \xi^\alpha, \quad \xi \in \mathbb{R}^d, \quad (2.5)$$

with $\xi^\alpha := \xi_1^{\alpha_1} \dots \xi_d^{\alpha_d}$, for all multi-index $\alpha \in \mathbb{N}^d$. The polynomial function $b : \mathbb{R}^d \rightarrow \mathbb{R}$ is called *the symbol* of the Cauchy problem (2.1).

As standard, we denote by \mathcal{S} the Schwartz space of rapidly decreasing functions on \mathbb{R}^d , and by \mathcal{S}' the corresponding dual space of tempered distributions. We recall that this space contains all (distributions represented by) polynomially growing functions.

Assumption 2.1. *There exists $T \in (0, \infty]$ such that $\hat{g}(t, \cdot) \in \mathcal{S}'$ for all $t \in [0, T)$.*

Under this assumption, we may introduce the so-called Green functions as the inverse Fourier transform with respect to the space variable:

$$g(t, \cdot) := \mathfrak{F}^{-1} \hat{g}(t, \cdot), \quad t \in [0, T), \quad (2.6)$$

in the distribution sense, i.e. $\langle \mathfrak{F}^{-1} \hat{g}(t, \cdot), \varphi \rangle = \langle \hat{g}(t, \cdot), \mathfrak{F}^{-1} \varphi \rangle$ for all $\varphi \in \mathcal{S}$, $t \in [0, T)$, where

$$\mathfrak{F}^{-1} \varphi(x) := (2\pi)^{-d/2} \int_{\mathbb{R}^d} e^{i\xi \cdot x} \varphi(\xi) d\xi, \quad \text{for all } \varphi \in \mathcal{S}.$$

Assumption 2.2. *For all $n = 1, \dots, N$:*

(i) $(t, x) \mapsto (g_n(t, \cdot) * \phi)(x)$ is continuous on $[0, T) \times \mathbb{R}^d$, for all bounded continuous function ϕ on \mathbb{R}^d ;

(ii) $g_n(t, \cdot)$ may be represented by a signed measure $g_n(t, dx) = g_n^+(t, dx) - g_n^-(t, dx)$, with total variation measure $|g_n| := g_n^+ + g_n^-$ absolutely continuous with respect to some probability measure μ_n ; the corresponding densities γ_n , γ_n^+ and γ_n^- , defined by

$$g_n^+(t, dx) = \gamma_n^+(t, x) \mu_n(t, dx), \quad g_n^-(t, dx) = \gamma_n^-(t, x) \mu_n(t, dx), \quad \gamma_n := \gamma_n^+ - \gamma_n^-,$$

satisfy $|\gamma_n(t, \cdot)|_\infty < \infty$, $\gamma_n(t, \mathbb{R}^d) < \infty$, and $\gamma_N(\cdot, \mathbb{R}^d) \in \mathbb{L}^1([0, t])$ for all $t \in [0, T)$.

The choice of the dominating measures $(\mu_n)_{1 \leq n \leq N}$ will be discussed in Example 2.10.

Proposition 2.3. *Let $F, (f_n)_{1 \leq n \leq N}$ be as in (2.3). Then, under Assumptions 2.1 and 2.2, the Cauchy problem (2.1)-(2.2) has a unique solution in $C_b^0([0, T) \times \mathbb{R}^d, \mathbb{R})$ given by*

$$u(t, x) := \sum_{n=1}^N p_n(f_n * g_n(t, \cdot))(x) + \int_0^t (F(t-s, \cdot) * g_N(s, \cdot))(x) ds, \quad t \in [0, T), \quad x \in \mathbb{R}^d. \quad (2.7)$$

We observe that the standard statement of the last proposition involves different assumptions on g_n , f_n and F . Namely, one may typically relax Assumption 2.2 and assume that f and F have bounded support so as to guarantee that the convolutions involved in the representation (2.7) are well-defined and satisfy the property $\mathfrak{F}(f_n * g_n(t, \cdot)) = (2\pi)^{\frac{d}{2}} \mathfrak{F}(f_n) \mathfrak{F}(g_n(t, \cdot))$ for all $t \in [0, T]$, $n = 1, \dots, N$. Our conditions in Proposition 2.3 are suitable for the subsequent use in the paper.

For the convenience of the reader, we report the proof of Proposition 2.3.

Proof. First, the conditions on the densities γ_n contained in Assumption 2.2 guarantee that the function u defined in (2.7) is bounded and continuous. Then the distribution represented by u is in \mathcal{S}' , and we may define the corresponding Fourier transform in the space variable $\hat{u}(t, \cdot) := \mathfrak{F}(u(t, \cdot))$ in the distribution sense. By standard calculation using the properties of the Fourier transform, we see that

$$\hat{u}(t, \xi) = \sum_{n=1}^N p_n \hat{f}_n(\xi) \hat{g}_n(t, \xi) + \int_0^t \hat{F}(t-s, \xi) \hat{g}_N(s, \xi) ds, \quad t \geq 0, \xi \in \mathbb{R}^d, \quad (2.8)$$

where $\hat{f} := \mathfrak{F}(f)$ and $\hat{F} := \mathfrak{F}(F(t, \cdot))$. By the definition of \hat{g} in (2.4), we see that, for every fixed $\xi \in \mathbb{R}^d$, the function $\hat{u}(\cdot, \xi)$ is the unique solution of the ODE

$$\sum_{n=1}^N a_n \partial_t^n \hat{u} - b(\xi) \hat{u} - \hat{F}(\cdot, \xi) = 0 \quad \text{on } \mathbb{R}_+, \quad (2.9)$$

$$\partial_t^{n-1} \hat{u}(0, \xi) = p_n \hat{f}_n(\xi), \quad n = 1, \dots, N, \quad (2.10)$$

which can be written equivalently as a first order ODE in terms of the function $\hat{v} := (\hat{u}, \partial_t \hat{u}, \dots, \partial_t^{N-1} \hat{u})^\top$:

$$\partial_t \hat{v} = B \hat{v} + \hat{F} e_N, \quad \text{on } \mathbb{R}_+, \quad \text{and} \quad \hat{v}(0, \cdot) = \sum_{n=1}^N p_n \hat{f}_n e_n.$$

Obviously, the last ODE has a unique solution with closed form obtained by the variation of the constant method $\hat{v}(t, \xi) := e^{tB(\xi)} \hat{v}(0, \xi) + \int_0^t e^{sB(\xi)} \hat{F}(t-s) e_N ds$, whose first entry induces the solution \hat{u} introduced in (2.8). To conclude the proof, it suffices to observe that any solution $\tilde{u} \in C_b^0([0, T] \times \mathbb{R}^d)$ of the Cauchy problem (2.1)-(2.2) has a well-defined Fourier transform in the distribution sense satisfying the ODE (2.9)-(2.10) for all fixed $\xi \in \mathbb{R}^d$. \square

Example 2.4 (Heat equation). *Let $N = 1$, $M = 2$, $b_\alpha = 0$ whenever $|\alpha| \leq 1$, and $b_{1,1} = b_{2,2} = 1$, $b_{1,2} = b_{2,1} = 0$. Then $B(\xi) = b(\xi) = -|\xi|^2$, $\xi \in \mathbb{R}^d$, and*

$$g_1(r, z) := (2\pi)^{-d} \int e^{-|\xi|^2 r + i\xi z} d\xi = (4\pi r)^{-d/2} e^{-\frac{|z|^2}{4r}}, \quad (r, z) \in \mathbb{R}_+ \times \mathbb{R}^d.$$

Example 2.5 (Airy equation). *Let $N = 1$, $M = 3$, $b_0 = 0$, $b_\alpha = 0$ whenever $|\alpha| \leq 2$, and $b_\alpha = 0$, and $\xi \mapsto b(\xi)$ odd in the sense that*

$$b(-\xi_j e_j + \sum_{\ell \neq j} \xi_\ell e_\ell) = -b_\alpha(\xi_j e_j + \sum_{\ell \neq j} \xi_\ell e_\ell) \quad \text{for all } j = 1, \dots, d.$$

Then $B(\xi) = b(\xi)$ is scalar valued, and

$$g_1(r, z) := (2\pi)^{-d} \int e^{i(-rb(\xi) + \xi z)} d\xi = \frac{1}{(3r)^{d/3}} \text{Ai}_b\left(\frac{z}{(3r)^{1/3}}\right), \quad (r, z) \in \mathbb{R}_+ \times \mathbb{R}^d,$$

where we introduced the d -dimensional Airy function

$$\text{Ai}_b(x) := (\pi)^{-d} \int_{\mathbb{R}_+^d} \cos\left(-\frac{b(\xi)}{3} + x \cdot \xi\right) d\xi, \quad x \in \mathbb{R}^d.$$

For $d = 1$, notice that $\int |g_1|(t, dz) = \infty$, so that Assumption 2.2 fails in this example.

Example 2.6 (Wave equation). Let $N = 2$, with $a_1 = 0$, $M = 2$, $b_\alpha = 0$ whenever $|\alpha| \leq 1$, and $b_{1,1} = b_{2,2} = 1$, $b_{1,2} = b_{2,1} = 0$. Then

$$B(\xi) = \begin{pmatrix} 0 & 1 \\ -|\xi|^2 & 0 \end{pmatrix}, \quad e^{B(\xi)r} = (2\pi)^{-\frac{d}{2}} \begin{pmatrix} \cos(r|\xi|) & |\xi|^{-1} \sin(r|\xi|) \\ -|\xi| \sin(r|\xi|) & \cos(r|\xi|) \end{pmatrix}, \quad \xi \in \mathbb{R}^d,$$

and, for $(r, z) \in \mathbb{R}_+ \times \mathbb{R}^d$,

$$g_2(r, dz) = (2\pi)^{-d} \int \frac{\sin(r|\xi|)}{|\xi|} e^{i\xi \cdot z} d\xi, \quad g_1(r, dz) = (2\pi)^{-d} \int \cos(r|\xi|) e^{i\xi \cdot z} d\xi = \partial_r g_2(r, dz).$$

We have (see Kirchhoff's formula in [10] for closed-expression for g_1 and g_2 in \mathbb{R}^d)

$$g_2(r, dz) = \begin{cases} \frac{1}{2} \mathbf{1}_{\{|z| < r\}} dz, & \text{for } d = 1 \\ \frac{1}{2\pi} (r^2 - |z|^2)^{-\frac{1}{2}} \mathbf{1}_{\{|z| < r\}} dz, & \text{for } d = 2 \\ \frac{\sigma_r(dz)}{4\pi r}, & \text{for } d = 3 \end{cases}$$

where $\sigma_r(dz)$ denotes the surface area on $\partial B(0, r)$. Assumption 2.2 are discussed in Example 2.10.

Example 2.7 (Beam equation). Let $d = 1$, $N = 2$, with $a_1 = 0$, $M = 4$, $b_1 = b_2 = b_3 = 0$, and $b_4 = 1$ corresponding to the fourth-order PDE $u_{tt} + \partial_x^4 u = 0$. Then,

$$g_1(r, z) = \partial_r g_2(r, z) \quad \text{and} \quad g_2(r, z) = \sqrt{r} G\left(\frac{z}{\sqrt{r}}\right),$$

where $G(0) = 1/\sqrt{2\pi}$, and

$$2G'(x) = \text{F}_s\left(\frac{x}{\sqrt{2\pi}}\right) - \text{F}_c\left(\frac{x}{\sqrt{2\pi}}\right), \quad x \in \mathbb{R}, \quad (2.11)$$

with the Fresnel integrals $\text{F}_c(x) := \int_0^x \cos(\pi t^2/2) dt$ and $\text{F}_s(x) := \int_0^x \sin(\pi t^2/2) dt$. Note that $\int_{\mathbb{R}} |G(x)| dx < \infty$ as $G(x) \underset{|x| \rightarrow \infty}{\sim} \frac{\sqrt{\frac{2}{\pi}} (\cos(\frac{x^2}{4}) - \sin(\frac{x^2}{4}))}{x^2}$.

Remark 2.8. In order to compute the Green functions g as the inverse Fourier transform of the associated \hat{g} , one needs to diagonalize the matrix $B(\xi)$. Direct examination reveals that the eigenvalues of the matrix $B(\xi)$ are the roots of the corresponding characteristic polynomial $b(\xi) - \sum_{n=1}^N a_n \lambda^n$. Assume that $B(\xi)$ has N distinct (simple) eigenvalues $(\lambda_j(\xi))_{1 \leq j \leq N} \in \mathbb{C}^N$. Then, denoting by $\text{diag}[\lambda]$ the diagonal matrix with $\text{diag}[\lambda]_{j,j} = \lambda_j$, it follows that

$$B(\xi) = P(\xi) \text{diag}[\lambda(\xi)] P(\xi)^{-1}, \quad \text{where } P(\xi)_{j,\ell} := \lambda_\ell(\xi)^j, \quad 1 \leq j, \ell \leq N.$$

The matrix $P(\xi)$ is the so-called Vandermonde matrix whose inverse is given by:

$$\{P(\xi)^{-1}\}_{j,n} = \frac{\Lambda_n(\xi)}{\lambda_j(\xi) \prod_{\ell \neq j} (\lambda_\ell - \lambda_j)(\xi)}, \quad j, n = 1, \dots, N,$$

where

$$\Lambda_N(\xi) = 1, \quad \text{and} \quad \Lambda_n(\xi) := (-1)^{n-1} \sum_{\substack{1 \leq \ell_1 \leq \dots \leq \ell_{N-n} \leq N \\ \ell_1, \dots, \ell_{N-n} \neq n}} (\lambda_{\ell_1} \cdots \lambda_{\ell_{N-n}})(\xi) \quad \text{for } n < N.$$

Therefore, by the definition of \hat{g}_n in (2.4), we have for $(r, \xi) \in \mathbb{R}_+ \times \mathbb{R}^d$:

$$\hat{g}_n(r, \xi) = (2\pi)^{-\frac{d}{2}} \Lambda_n(\xi) \sum_{j=1}^N \frac{e^{r \lambda_j(\xi)}}{\prod_{\ell \neq j} (\lambda_\ell - \lambda_j)(\xi)}, \quad n = 1, \dots, N. \quad (2.12)$$

□

Remark 2.9. All of the results of the present paper extend to the case of Cauchy problems for complex-valued functions, with coefficients $(a_n)_{1 \leq n \leq N}$ and $(b_\alpha)_{\alpha \in \mathbb{N}_M^d}$ in \mathbb{C} , thus allowing to include, for instance, the Schrödinger equation (see Example 2.11) and its semilinear extension as the Gross–Pitaevskii PDE (see Section 5.4). This extension follows by simply applying the methodology described throughout the paper separately to the real part and the imaginary part of the representation. □

2.3 Probabilistic representation

Let $(\Omega, \mathcal{F}, \mathbb{P})$ be a probability space supporting two random variables τ and I , with

$$\tau \text{ and } I \text{ independent, } \mathbb{P}[\tau \in dt] = \rho(t) \mathbf{1}_{\{t \geq 0\}} dt \text{ and } \mathbb{P}[I = n] = p_n, \quad n = 1, \dots, N, \quad (2.13)$$

for some $C^0(\mathbb{R}_+, \mathbb{R})$ density function $\rho > 0$ on $(0, \infty)$. We shall denote $\bar{\rho}(t) := \int_t^\infty \rho(s) ds$.

Recall the densities γ_n and the dominating probability measures $\mu_n(t, \cdot)$, $n = 1, \dots, N$, introduced in Assumption 2.2. For all $t \geq 0$, we introduce the random variables

$$X_t^n := x + Z_t^n \text{ independent of } (I, \tau), \text{ with } \mathbb{P}[Z_t^n \in dz] = \mu_n(t, dz), \quad n = 1, \dots, N. \quad (2.14)$$

Example 2.10. As $g_n(t, \cdot) \in \mathbb{L}^1(\mathbb{R}^d)$ by Assumption 2.2, we may choose $\mu_n(t, dz) = \|g_n(t, \cdot)\|_{\mathbb{L}^1(\mathbb{R}^d)}^{-1} |g_n|(t, dz)$ and $\gamma_n(t, z) = \text{sgn}(g_n)(t, z) \|g_n(t, \cdot)\|_{\mathbb{L}^1(\mathbb{R}^d)}$, where $\text{sgn}(\alpha) := \mathbf{1}_{\{\alpha \geq 0\}} - \mathbf{1}_{\{\alpha < 0\}}$.

(i) For the heat equation, we directly compute that $\|g_1(t, \cdot)\|_{\mathbb{L}^1(\mathbb{R}^d)} = 1$. Then, choosing $\gamma_1(t, \cdot) = 1$ leads to $Z_t^1 = \sqrt{2t}Z$ with $Z \in N(0, I_d)$ a d -Gaussian random variable.

(ii) For the wave equation, we directly compute that $\|g_2(t, \cdot)\|_{\mathbb{L}^1(\mathbb{R}^d)} = t$. Then, choosing $\gamma_2(t, \cdot) = t$ leads to $Z_t^2 = tZ$ where

- in dimension $d = 1$, Z has a uniform distribution on $[-1, 1]$
- in dimension $d = 2$, the law of Z is defined by the density $\frac{1}{2\pi} \frac{1}{\sqrt{1-z^2}} \mathbf{1}_{|z| < 1}$,
- in dimension $d = 3$, the law of Z is $\frac{1}{4\pi} \mu_{S^2}(dz)$, where μ_{S^2} denotes the volume measure on the unit sphere.

(iii) For the Beam equation, $\gamma_2(t, z) = t \text{sgn}(G(z)) \|G\|_{\mathbb{L}^1(\mathbb{R}^d)}$ where G is defined in (2.11). We have $Z_t^2 = \sqrt{t}Z$ with Z distributed according to $\|G\|_{\mathbb{L}^1(\mathbb{R}^d)}^{-1} |G(z)|$.

Example 2.11 (Schrödinger equation and analytical continuation). *The Duhamel formula for the Schrödinger equation of a free particle (with source term)*

$$i\partial_t u = -\frac{1}{2}\Delta u + F, \quad u(0, x) = f_1(x) \quad (2.15)$$

is

$$u(t, x) = \int_{\mathbb{R}^d} dy f_1(y) g_1(t, x-y) - i \int_0^t \int_{\mathbb{R}^d} dy g_1(s, x-y) F(s, y) ds$$

with $g_1(t, x) = \frac{e^{-\frac{x^2}{2it}}}{(2\pi it)^{\frac{d}{2}}}$. Note the coefficient $-i$ in front of F as $a_1 = i$ here. By setting $y - x = \sqrt{it}z$, this can be written as

$$u(t, x) = \int_{-\infty e^{-\frac{i\pi}{4}}}^{\infty e^{-\frac{i\pi}{4}}} dz f_1(x + e^{\frac{i\pi}{4}} \sqrt{t}z) e^{-\frac{z^2}{2}} - i \int_0^t ds \int_{-\infty e^{-\frac{i\pi}{4}}}^{\infty e^{-\frac{i\pi}{4}}} F(s, x + e^{\frac{i\pi}{4}} \sqrt{s}z) e^{-\frac{z^2}{2}} dz ds$$

By assuming that $|f_1(x + e^{\frac{i\pi}{4} + i\theta} \sqrt{t}R)| e^{-\frac{R^2}{2} \cos(2\theta)}$ and $|F(s, x + e^{\frac{i\pi}{4} + i\theta} \sqrt{s}R)| e^{-\frac{R^2}{2} \cos(2\theta)}$ goes to zero as $R \rightarrow \infty$ when $\theta \in [-\frac{\pi}{4}, 0]^d \cup [\frac{3\pi}{4}, \pi]^d$, the integration over z in $[-\infty e^{-\frac{i\pi}{4}}, \infty e^{-\frac{i\pi}{4}}]^d$ can be deformed into $[-\infty, +\infty]^d$ by analytical continuation and we obtain

$$u(t, x) = \mathbb{E}[f_1(x + e^{\frac{i\pi}{4}} \sqrt{t}Z)] - i \int_0^t ds \mathbb{E}[F(s, x + e^{\frac{i\pi}{4}} \sqrt{s}Z)] ds$$

with $Z \in N(0, I_d)$.

The following representation result is a simple rewriting of Proposition 2.3 in terms of the last notations.

Proposition 2.12. *Let $(f_n)_{1 \leq n \leq N}$ and F be as in (2.3). Then, under Assumptions 2.1 and 2.2, the unique $C_b^0([0, T] \times \mathbb{R}^d, \mathbb{R})$ solution of the Cauchy problem (2.1)-(2.2) is:*

$$u(t, x) = \mathbb{E} \left[\mathbf{1}_{\{\tau \geq t\}} \frac{\gamma_I(t, Z_t^I)}{\bar{\rho}(t)} f_I(X_\tau^I) + \mathbf{1}_{\{\tau < t\}} \frac{\gamma_N(\tau, Z_\tau^N)}{\rho(\tau)} F(t - \tau, X_\tau^N) \right], \quad t < T, x \in \mathbb{R}^d.$$

3 A first class of semilinear Cauchy problems

In this section, we consider the semilinear Cauchy problem:

$$\sum_{n=1}^N a_n \partial_t^n u - \sum_{\alpha \in \mathbb{N}_M^d} b_\alpha D^\alpha u - \sum_{j \geq 0} q_j c_j u^j = 0 \quad \text{on } \mathbb{R}_+ \times \mathbb{R}^d, \quad (3.1)$$

$$\partial_t^{n-1} u(0, \cdot) = p_n f_n \quad \text{on } \mathbb{R}^d, \quad n = 1, \dots, N, \quad (3.2)$$

where the nonlinearity is defined by means of an atomic probability measure $(q_j)_{j \geq 0}$, with $q_j \geq 0$ and $\sum_{j \geq 0} q_j = 1$, together with the functions

$$c_j : \mathbb{R}_+ \times \mathbb{R}^d \longrightarrow \mathbb{R}, \quad j \geq 0, \quad \text{bounded continuous, for all } j \geq 0. \quad (3.3)$$

In order for the power series on the right hand-side of (3.1) to be well defined, we assume that

$$H_1(s) := \sum_{j \geq 0} q_j \|c_j\|_\infty s^j. \quad (3.4)$$

has a strictly positive radius of convergence, so that it is well defined at least in some neighborhood of the origin.

3.1 The branching mechanism

A particle of generation $\nu \in \mathbb{N}$ is a multi-integer $k := (0, k_1, \dots, k_\nu) \in \mathbf{K}_\nu := \{0\} \times \mathbb{N}^\nu$. We set by convention $\mathbf{K}_0 := \{0\}$, and we denote by $\mathbf{K} := \cup_{\nu \geq 0} \mathbf{K}_\nu$ the collection of all particles. For $\nu \geq 1$, and a particle $k := (0, k_1, \dots, k_\nu) \in \mathbf{K}_\nu$, we denote by k_- its parent particle

$$k_- := (0, k_1, \dots, k_{\nu-1}) \in \mathbf{K}_{\nu-1}.$$

We next introduce independent families of random variables $(\tau_k)_{k \in \mathbf{K}}$, $(I_k^0)_{k \in \mathbf{K}}$, and $(J_k)_{k \in \mathbf{K}}$:

- I_k^0 and τ_k are iid copies of the random variables I and τ , respectively, as introduced in (2.13);
- J_k are iid random variables with $\mathbb{P}[J_k = j] = q_j$ for all $j \geq 0$.

The time occurrences of the branching events, prior to t , are recorded through the sequence $(T_k^t)_k$ defined for all $t \geq 0$ by:

$$T_{0-}^t := 0, \quad \text{and} \quad T_k^t := t \wedge (T_{k-}^t + \tau_k), \quad \text{for all } k \in \mathbf{K}.$$

With these notations, each particle k lives on the time interval $[T_{k-}^t, T_k^t]$. The branching mechanism is then the following:

- Start from particle 0 at the time origin.
- At the first branching time $T_0^t = t \wedge \tau_0$, particle zero dies out; if $T_0^t < t$, it generates J_0 descendants labelled $(0, 1), \dots, (0, J_0)$.
- Each descendant particle k undergoes the same behavior, independently of its peer-particles: it dies out at the branching time T_k^t , and generates J_k descendants labelled $(k, 1), \dots, (k, J_k)$, whenever $T_k^t < t$.
- We denote by \mathcal{K}_s^t the collection of all living particles at time s , and by $\overline{\mathcal{K}}_t := \cup_{s \leq t} \mathcal{K}_s^t$ the collection of all particles which have been living prior to time t . For simplicity we set $\mathcal{K}_t := \mathcal{K}_t^t$.
- We finally denote for all particle $k \in \overline{\mathcal{K}}_t$:

$$I_k^t := I_k^0 \mathbf{1}_{k \in \mathcal{K}_t} + N \mathbf{1}_{k \in \overline{\mathcal{K}}_t \setminus \mathcal{K}_t}. \quad (3.5)$$

3.2 Probabilistic representation

Given the branching mechanism defined in the previous subsection, we introduce the corresponding branching process:

$$X_0^0 = x, \quad \text{and} \quad X_{T_k^t}^k := X_{T_{k-}^t}^{k-} + Z_{T_k^t}^k, \quad \text{for all } k \in \overline{\mathcal{K}}_t, \quad (3.6)$$

where the distribution of $Z_{T_k^t}^k$ depends on the type of the particle:

$$\mathbb{P}[Z_{T_k^t}^k \in dz | I_k^t, \Delta T_k^t] = \mu_{I_k^t}(\Delta T_k^t, dz), \quad \text{with} \quad \Delta T_k^t := T_k^t - T_{k-}^t. \quad (3.7)$$

We now introduce, for all $t \geq 0$, and $x \in \mathbb{R}^d$, the random variable

$$\xi_{t,x} := \prod_{k \in \mathcal{K}_t} \frac{\gamma_{I_k^t}(\Delta T_k^t, Z_{T_k^t}^k)}{\bar{\rho}(\Delta T_k^t)} f_{I_k^t}(X_{T_k^t}^k) \prod_{k \in \overline{\mathcal{K}}_t \setminus \mathcal{K}_t} \frac{\gamma_N(\Delta T_k^t, Z_{T_k^t}^k)}{\rho(\Delta T_k^t)} c_{J_k}(t - T_k^t, X_{T_k^t}^k). \quad (3.8)$$

Recall the power series H_1 defined in (3.4).

Assumption 3.1. *The power series H_1 has a radius of convergence $R_1 \in (0, \infty]$, and*

- $r_1 := \sup_{1 \leq n \leq N} \|f_n\|_\infty \|\gamma_n\|_\infty < R_1$;
- there are constants $T > 0$ and $s_1 > r_1$ such that $\int_{r_1}^{s_1} H_1(s)^{-1} ds = T \|\gamma_N\|_\infty$.

Theorem 3.2. *Let $f \in C_b^0(\mathbb{R}^d, \mathbb{R}^n)$ and $c_j \in C_b^0(\mathbb{R}_+ \times \mathbb{R}^d, \mathbb{R})$. Then, under Assumptions 2.1, 2.2 and 3.1, we have $\xi_{t,x} \in \mathbb{L}^1$ and $u(t,x) := \mathbb{E}[\xi_{t,x}]$ is the unique solution of the semilinear Cauchy problem (3.1)-(3.2).*

Proof. 1. We first verify the \mathbb{L}^1 -integrability of $\xi_{t,x}$. From the expression of $\xi_{t,x}$ in (3.8), we see that

$$|\xi_{t,x}| \leq \chi_t := \prod_{k \in \mathcal{K}_t} \frac{r_1}{\bar{\rho}(\Delta T_k^t)} \prod_{k \in \bar{\mathcal{K}}_t \setminus \mathcal{K}_t} \frac{\bar{c}_{J_k} \|\gamma_N\|_\infty}{\rho(\Delta T_k^t)}. \quad (3.9)$$

Our objective is now to prove that $t \mapsto \mathbb{E}\chi_t < \infty$. By standard arguments, this function is related to the ODE

$$\partial_t w = \|\gamma_N\|_\infty H_1(w) \quad \text{and} \quad w(0) = r_1. \quad (3.10)$$

Let us first verify that this ODE has a non-exploding solution under the second condition in Assumption 3.1. Since the power series function H_1 has radius R_1 , the condition $r_1 < R_1$ is necessary to find a (finite) solution of the last ODE. Next, for an arbitrary $L > 0$, notice that H_1 is Lipschitz on $[-L, L]$, so that the last ODE has a unique non-negative solution w up to T_L , where

$$\lim_{t \rightarrow T_L} w(t) = L \quad \text{whenever} \quad T_L < \infty,$$

and the non-negativity of w follows from the fact that $w(0) > 0$ and all derivatives $H_1^{(i)}(0) \geq 0$ for all $i \geq 0$. Then, it follows from direct integration of the ODE that

$$\|\gamma_N\|_\infty T_L = \int_0^{T_L} \frac{\partial_s w}{H_1(w)}(s) ds = \int_{r_1}^{L} \frac{ds}{H_1(s)} \quad \text{as long as} \quad T_L < \infty.$$

This proves that w is non-exploding solution on $[0, T]$ if and only if $\int_{r_1}^{s_1} \frac{ds}{H_1(s)} = \|\gamma_N\|_\infty T$ for some constant $s_1 > r_1$.

Under the last condition, we obtain by direct integration of the ODE (3.10) that:

$$\begin{aligned} w(t) &= r_1 + \|\gamma_N\|_\infty \sum_{j \geq 0} q_j \bar{c}_j \int_0^t w(s)^j ds \\ &= \mathbb{E} \left[\frac{r_1 \mathbf{1}_{\{\tau_0 \geq t\}}}{\bar{\rho}(t)} + \mathbf{1}_{\{\tau_0 < t\}} \frac{\|\gamma_N\|_\infty \|c_{J_0}\|_\infty}{\rho(\tau_0)} w(\tau_0)^{J_0} \right] \\ &= \mathbb{E} \left[\frac{r_1 \mathbf{1}_{\{\tau_0 \geq t\}}}{\bar{\rho}(t)} + \mathbf{1}_{\{\tau_0 < t\}} \frac{\|\gamma_N\|_\infty \|c_{J_0}\|_\infty}{\rho(\tau_0)} \right. \\ &\quad \left. \prod_{k \in \mathcal{K}_t^1} \left(\frac{r_1 \mathbf{1}_{\{T_k^t \geq t\}}}{\bar{\rho}(\tau_k)} + \mathbf{1}_{\{T_k^t < t\}} \frac{\|\gamma_N\|_\infty \|c_{J_0}\|_\infty}{\rho(\tau_k)} w(T_k^t)^{J_k} \right) \right], \end{aligned}$$

where \mathcal{K}_t^n denotes the particles generated after n defaults prior to t and the last two equalities follow from the definition of the branching mechanism together with the tower property. Iterating up to the n -th generation, this provides:

$$v(t) = \mathbb{E}\chi_t^n, \quad \text{where} \quad \chi_t^n := \prod_{k \in \mathcal{K}_t^n} \frac{r_1}{\bar{\rho}(\Delta T_k^t)} \prod_{k \in \cup_{j < n} \mathcal{K}_t^j} \frac{\|\gamma_N\|_\infty \|c_{J_0}\|_\infty}{\rho(\Delta T_k^t)} \\ + \prod_{k \in \cup_{j \leq n} \mathcal{K}_t^j} \frac{\|\gamma_N\|_\infty \|c_{J_0}\|_\infty}{\rho(\Delta T_k^t)} \prod_{k \in \mathcal{K}_t^{n+1}} w(T_{k-}^t)^{J_{k-}}.$$

As $w \geq 0$, it follows that $\chi_t^n \geq 0$, and we then deduce from Fatou's lemma that

$$w(t) = \liminf_{n \rightarrow \infty} \mathbb{E}\chi_t^n \geq \mathbb{E}\chi_t.$$

Since $w(t) < \infty$, this provides the required integrability of the bound χ_t in (3.9).

2. We next prove that $v(t, x) := \mathbb{E}[\xi_{t,x}]$ is the unique solution of the nonlinear Cauchy problem (3.1). To see this, observe that

$$\xi_{t,x} = \mathbf{1}_{\{\tau_0 \geq t\}} \frac{\gamma_{I_0^t}(t, Z_t^0)}{\bar{\rho}(t)} f_{I_0^t}(X_t^0) + \mathbf{1}_{\{\tau_0 < t\}} \frac{\gamma_N(\tau_0, Z_{\tau_0}^0)}{\rho(\tau_0)} c_{J_0}(t - \tau_0, X_{\tau_0}^0) \prod_{j=1}^{J_0} \xi_{t-\tau_0, X_{\tau_0}^0}^{(j)},$$

where $\xi_{t-\tau_0, X_{\tau_0}^0}^{(j)}$ have the same distribution, conditional on $(\tau_0, X_{\tau_0}^0)$. Taking expectations, and using the tower property, it follows from the definition of the function v that

$$v(t, x) = \mathbb{E} \left[\mathbf{1}_{\{\tau_0 \geq t\}} \frac{\gamma_{I_0^t}(t, Z_t^0)}{\bar{\rho}(t)} f_{I_0^t}(X_t^0) + \mathbf{1}_{\{\tau_0 < t\}} \frac{\gamma_N(\tau_0, Z_{\tau_0}^0)}{\rho(\tau_0)} c_{J_0}(t - \tau_0, X_{\tau_0}^0) v(t - \tau_0, X_{\tau_0}^0)^{J_0} \right] \\ = \sum_{n=1}^N \int p_n(f_n * g_n(t, \cdot))(x) + \int_0^t (F(t-s, \cdot) * g_N(s, \cdot))(x) ds,$$

where $F := \sum_{j \geq 0} q_j c_j v^j$. Since v is bounded, it follows from Proposition 2.3 that v is the unique $C_b^0([0, T] \times \mathbb{R}^d, \mathbb{R})$ solution of the nonlinear Cauchy problem (3.1)-(3.2). \square

Remark 3.3 (Finite propagation speed). From the simulation of Z_t^2 in the case of the wave equation, we deduce directly the finite propagation speed property: If $f_1 = f_2 = 0$ on a ball $B(x_0, t_0)$ of center x_0 and radius t_0 then $u = 0$ within the cone $K(x_0, t_0) := \{(x, t) : 0 \leq t \leq t_0 \text{ and } |x - x_0| \leq t_0 - t\}$. \square

The next statement provides some sufficient conditions for $\xi_{t,x}$ to have a finite p -th moment. In particular, for $p = 2$, this guarantees that the error estimate of the Monte Carlo approximation induced by the representation of Theorem 3.2 is characterized by the standard central limit theorem, and is in particular independent of the dimension d of the space variable x . Our first requirement is stated on the power series:

$$H_p(s) := \sum_{j \geq 0} (q_j \|c_j\|_\infty)^p s^j. \quad (3.11)$$

Assumption 3.4. The power series H_p has a radius of convergence $R_p \in (0, \infty]$, and

- (i) $r_p := \max_{1 \leq n \leq N} \|f_n\|_\infty^p \|\gamma_n \bar{\rho}^{\frac{1-p}{p}}\|_\infty^p < R_p$, and $\alpha_p := \|\gamma_N \rho^{\frac{1-p}{p}}\|_\infty^p < \infty$;
- (ii) there are constants $T > 0$ and $s_p > r_p$ such that $\int_{r_p}^{s_p} H_p(s)^{-1} ds = \alpha_p T$.

Theorem 3.5. Let $f \in C_b^0(\mathbb{R}^d, \mathbb{R}^n)$ and $c_j \in C_b^0(\mathbb{R}_+ \times \mathbb{R}^d, \mathbb{R})$. Then, under Assumptions 2.1, 2.2 and 3.4, we have $\xi_{t,x} \in \mathbb{L}^p$ for all $t \in [0, T]$.

Proof. Similar to the calculation in the previous proof, we have

$$|\xi_{t,x}|^p \leq \prod_{k \in \mathcal{K}_t} \frac{r_p}{\bar{\rho}(\Delta T_k^t)} \prod_{k \in \bar{\mathcal{K}}_t \setminus \mathcal{K}_t} \frac{\alpha_p \|c_{J_k}\|_\infty^p}{\rho(\Delta T_k^t)}.$$

The right hand-side is now related to the ODE

$$\partial_t w = \alpha_p H_p(w) \quad \text{and} \quad w(0) = r_p.$$

the integrability of $|\xi_{t,x}|^p$ can now be verified by following the same line of argument as in Step 1 of the proof of Theorem 3.2. \square

4 Further nonlinear Cauchy problems

In this section, we consider the following semilinear Cauchy problem with polynomial nonlinearity in the pair (u, Du) :

$$\sum_{n=1}^N a_n \partial_t^n u - \sum_{\alpha \in \mathbb{N}_M^d} b_\alpha D^\alpha u - \sum_{j \geq 0} q_j c_{j,0} u^{\ell_{j,0}} \prod_{h=1}^H (c_{j,h} \cdot Du)^{\ell_{j,h}} = 0 \quad \text{on } \mathbb{R}_+ \times \mathbb{R}^d \quad (4.1)$$

$$\partial_t^{n-1} u(0, \cdot) = p_n f_n \quad \text{on } \mathbb{R}^d, \quad n = 1, \dots, N, \quad (4.2)$$

where $(\ell_j)_{j \geq 0} \subset \mathbb{N}^{1+H}$ is a sequence of vector integers $\ell_j = (\ell_{j,0}, \dots, \ell_{j,H})$, and the nonlinearity is defined by means of an atomic probability measure $(q_j)_{j \geq 0}$, with $q_j > 0$ and $\sum_{j \geq 0} q_j = 1$, together with the functions

$$c_{j,0} : \mathbb{R}_+ \times \mathbb{R}^d \longrightarrow \mathbb{R}, \quad \text{bounded continuous, and } c_{j,h} : \mathbb{R}_+ \times \mathbb{R}^d \longrightarrow B_1(\mathbb{R}^d), \quad j \geq 0, \quad 1 \leq h \leq H. \quad (4.3)$$

Here, $B_1(\mathbb{R}^d)$ denotes the unit ball in \mathbb{R}^d in the sense of the Euclidean norm. Notice that the above Cauchy problem covers the particular case of (3.1)-(3.2) by setting $\ell_{j,h} = 0$ for all $h = 1, \dots, H$.

Recall the Green functions $g = (g_n)_{1 \leq n \leq N}$ from (2.6). By standard Fourier transform theory, we have

$$\partial_x g_n(t, \cdot) = i \mathfrak{F}^{-1}(\xi \mapsto \xi \hat{g}_n(t, \xi)), \quad t \geq 0, \quad n = 1, \dots, N.$$

Similar to the probability measures μ_n introduced in Assumption 2.2, we now assume that the distributions $\partial_x g_n$ can be represented by signed measures for which we may introduce dominating measures for $|\partial_x g_n(t, \cdot)| := \sum_{m=1}^d |\partial_{x_m} g_n(t, \cdot)|$.

Assumption 4.1. (i) For all $n = 1, \dots, N$ and $m = 1, \dots, d$, the distribution $\partial_{x_m} g_n(t, \cdot)$ may be represented by a signed measure with total variation $|\partial_{x_m} g_n(t, \cdot)|$.
(ii) The measure $|\partial_x g_n(t, \cdot)|$ is absolutely continuous with respect to some probability measure $\mu_n^1(t, \cdot)$ so that we may define the density vector $\gamma_n^1 = (\gamma_{n,1}^1, \dots, \gamma_{n,d}^1)$ by

$$\partial_x g_n(t, dx) = \gamma_n^1(t, x) \mu_n^1(t, dx), \quad t \geq 0, \quad x \in \mathbb{R}^d.$$

Notice that although $\gamma_n^1(t, \cdot)$ is defined $d\mu_n^1(t, \cdot)$ -a.s. this will be sufficient for our needs. Our starting point is the following ‘‘automatic differentiation property’’, which follows by direct differentiation of the Duhamel formula of Proposition 2.3.

Proposition 4.2. In addition to the conditions of Proposition 2.3, let Assumption 4.1 hold true. Then, the solution u of the linear Cauchy problem (2.1)-(2.2) is differentiable with respect to the space variable with

$$\partial_x u(t, x) = \mathbb{E} \left[\mathbf{1}_{\{\tau \geq t\}} \frac{\gamma_I^1(t, Z_t^{1,I})}{\bar{\rho}(t)} f_I(X_\tau^{1,I}) + \mathbf{1}_{\{\tau < t\}} \frac{\gamma_N^1(\tau, Z_\tau^{1,N})}{\rho(\tau)} F(t - \tau, X_\tau^{1,N}) \right],$$

where τ is a random time with density ρ , and $X_\tau^{1,I} = x + Z_\tau^{1,I}$, with $Z_t^{1,I}$ distributed as $\mu^1(t, \cdot)$ independent of τ .

Example 4.3. Let us illustrate the last result on our main examples.

- (i) Heat equation: we directly compute that $\partial_x g_1(t, dx) = -\frac{x}{2t} g_1(t, dx)$.
- (ii) Wave equation, $d = 1$: we have $\partial_x g_2(t, dx) = \frac{1}{2} (\delta(x+t) - \delta(x-t)) dx$, but $\partial_x g_1 = \partial_{xx}^2 g_2$ can not be represented as a signed measure, thus violating Assumption 4.1. However, we may still handle the one dimensional wave equation by reducing to the case $f_1 = 0$, see Section 5.2.
- (iii) Wave equation, $d > 1$: Assumption 4.1 is not satisfied as $\partial_x g_1$ and $\partial_x g_2$ involve first-order derivative of the delta function supported on the light cone $\{(x, t) : t^2 - x^2 = 0\}$.
- (iv) Beam equation: we have $\partial_x g_2(t, dx) = G'(\frac{x}{\sqrt{t}}) dx$.

In order to introduce the probabilistic representation of the solution of (4.1)-(4.2), we consider the branching mechanism defined in Subsection 3.1, where we modify the definition of the independent iid random variables $(J_k)_{k \in \mathbf{K}}$ and we introduce the types of particles $(\theta_k^t)_{k \in \mathbf{K}}$ as follows:

- $\mathbb{P}[J_k = \ell_j] = q_j$ for all $j \geq 0$, and we denote $\bar{J}_{k,-1} := 0$, $\bar{J}_{k,h} := \bar{J}_{k,h-1} + J_{k,h}$, $h = 0, \dots, H$;
- an arbitrary particle $k \in \bar{\mathcal{K}}_t \setminus \mathcal{K}_t$ branches at time T_k into $\bar{J}_{k,H}$ particles; the h -th block of descendent particles are labelled

$$(k, j), \quad j = \bar{J}_{k,h-1} + 1, \dots, \bar{J}_{k,h} \quad h = 0, \dots, H; \tag{4.4}$$

- we assign to the h -th block of new particles (4.4) the types:

$$\theta_{(0)}^t := 0 \text{ and } \theta_{(k,j)}^t := h, \text{ for } k \in \bar{\mathcal{K}}_t \setminus \mathcal{K}_t, j = \bar{J}_{k,h-1} + 1, \dots, \bar{J}_{k,h}, h = 0, \dots, H. \quad (4.5)$$

Finally, we introduce a branching process which differs slightly from (3.6)-(3.7). Let

$$\hat{X}_0^0 = x, \quad \text{and} \quad \hat{X}_{T_k^t}^k := \hat{X}_{T_{k-}^t}^{k-} + \hat{Z}_{T_k^t}^k, \quad \text{for all } k \in \bar{\mathcal{K}}_t, \quad (4.6)$$

where the distribution of $\hat{Z}_{T_k^t}^k$ depends on the type of the particle:

$$\mathbb{P}[\hat{Z}_{T_k^t}^k \in dz | I_k^t, \theta_k^t, \Delta T_k^t] = \mathbf{1}_{\{\theta_k^t=0\}} \mu_{I_k^t}(\Delta T_k^t, dz) + \mathbf{1}_{\{\theta_k^t \neq 0\}} \mu_{I_k^t}^1(\Delta T_k^t, dz), \quad (4.7)$$

with $\Delta T_k^t := T_k^t - T_{k-}^t$. The main goal of this section is to provide a representation of the solution of the Cauchy problem (4.1)-(4.2) by means of the random variable

$$\hat{\xi}_{t,x} := \prod_{k \in \mathcal{K}_t} \frac{\mathcal{W}_k}{\bar{\rho}(\Delta T_k^t)} [f_{I_k^t}(\hat{X}_t^k) - \mathbf{1}_{\{\theta_k^t \neq 0\}} f_{I_k^t}(\hat{X}_{T_{k-}^t}^k)] \prod_{k \in \bar{\mathcal{K}}_t \setminus \mathcal{K}_t} \frac{\mathcal{W}_k}{\rho(\Delta T_k^t)} c_{J_k,0}(t - T_k^t, \hat{X}_{T_k^t}^k), \quad (4.8)$$

where the random weights \mathcal{W}_k are given by

$$\mathcal{W}_k := \mathbf{1}_{\{\theta_k^t=0\}} \gamma_{I_k^t}(\Delta T_k^t, \hat{Z}_{T_k^t}^k) + \mathbf{1}_{\{\theta_k^t \neq 0\}} c_{J_k, \theta_k^t}(t - T_{k-}^t, \hat{X}_{T_{k-}^t}^{k-}) \cdot \gamma_{I_k^t}^1(\Delta T_k^t, \hat{Z}_{T_k^t}^k), \quad k \in \bar{\mathcal{K}}_t.$$

Recall the power series H_p introduced in (3.11), and define

$$\hat{r}_p := \sup_{\substack{0 \leq t \leq T \\ 1 \leq n \leq N}} \frac{\{\|f_n\|_\infty^p \int |\gamma_n(t, z)|^{p-1} |g_n(t, dz)|\} \vee \{\|\nabla f_n\|_\infty^p \int |z|^p |\gamma_n^1(t, z)|^{p-1} |\partial_x g_n(t, dz)|\}}{\bar{\rho}(t)^{p-1}} \quad (4.9)$$

$$\hat{\alpha}_p := \sup_{0 \leq t \leq T} \frac{\{\int |\gamma_N(t, z)|^{p-1} |g_N(t, dz)|\} \vee \{\int |\gamma_N^1(t, z)|^{p-1} |\partial_x g_N(t, dz)|\}}{\rho(t)^{p-1}}. \quad (4.10)$$

The following result provides a probabilistic representation of the solution of the semilinear Cauchy (4.1) under a condition involving the above \hat{r}_p and $\hat{\alpha}_p$, that will be guaranteed in Proposition 4.9 below to hold under some sufficient conditions.

Theorem 4.4. *Let Assumptions 2.1, 2.2, and 4.1 hold true, and assume that f is bounded and Lipschitz. Let $p > 1$, and ρ be a positive density function with support on $(0, \infty)$ such that the constants \hat{r}_p and $\hat{\alpha}_p$ defined in (4.9)-(4.10) satisfy:*

$$\hat{r}_p < R_p, \quad \hat{\alpha}_p < \infty \quad \text{and} \quad \int_{\hat{r}_p}^{\hat{s}_p} \frac{ds}{H_p(s)} = \hat{\alpha}_p T \text{ for some } \hat{s}_p > \hat{r}_p \text{ and } T > 0. \quad (4.11)$$

Then, $\hat{\xi}_{t,x} \in \mathbb{L}^p$ for $t \in [0, T]$, and $u(t, x) := \mathbb{E}[\hat{\xi}_{t,x}]$ is the unique bounded continuous solution of the semilinear Cauchy problem (4.1).

Proof. Similar to the proof of Theorem 3.2, we directly estimate that

$$\begin{aligned} |\hat{\xi}_{t,x}|^p &\leq \prod_{k \in \mathcal{K}_t} \frac{1}{\bar{\rho}(\Delta T_k^t)^p} \left(\mathbf{1}_{\theta_k^t=0} \|f_{I_k^t}\|_\infty^p |\gamma_{I_k^t}(\Delta T_k^t, \hat{Z}_{T_k^t}^k)|^p + \mathbf{1}_{\theta_k^t \neq 0} \|\nabla f_{I_k^t}\|_\infty^p |Z_{T_k^t}^k|^p |\mathcal{W}_k|^p \right) \\ &\quad \times \prod_{k \in \bar{\mathcal{K}}_t \setminus \mathcal{K}_t} \left| \frac{\mathcal{W}_k}{\rho(\Delta T_k^t)} \right|^p \|c_{J_k,0}\|_\infty^p. \end{aligned}$$

By the independence of the ΔT_k^t 's and the $Z_{T_k^t}^k$'s, this provides

$$\mathbb{E} |\hat{\xi}_{t,x}|^p \leq \mathbb{E} \prod_{k \in \mathcal{K}_t} \frac{\hat{r}_p}{\bar{\rho}(\Delta T_k^t)} \prod_{k \in \bar{\mathcal{K}}_t \setminus \mathcal{K}_t} \frac{\hat{\alpha}_p \|c_{J_k,0}\|_\infty^p}{\rho(\Delta T_k^t)}.$$

The required result follows by the same line of argument as in the proof of Theorem 3.2. \square

In the rest of this section, we provide sufficient conditions which guarantee that the constants \hat{r}_p and $\hat{\alpha}_p$, involved in Condition (4.11), are finite. In preparation for this, we need some estimates on the Green functions g_n , $n = 1, \dots, N$. Recall that $\{\alpha \in \mathbb{N}_M^d : |\alpha| = M \text{ and } b_\alpha \neq 0\} \neq \emptyset$, and define *the principal symbol*

$$b_M(\xi) := i^M \sum_{|\alpha|=M} b_\alpha \xi^\alpha = e^{i\pi \eta_M(\xi)} |b_M(\xi)|, \text{ where } \eta_M(\xi) := \frac{1 + M - \text{sg}\{b_M(\xi)\}}{2}; \xi \in \mathbb{R}^d.$$

Lemma 4.5. *For all $\xi \in \mathbb{R}^d$, the matrix $B(\varepsilon^{-1}\xi)$ has N simple eigenvalues $\lambda_n(\varepsilon^{-1}\xi)$, $n = 1, \dots, N$, for sufficiently small $\varepsilon > 0$, with asymptotics*

$$\lim_{\varepsilon \searrow 0} \varepsilon^{\frac{1}{\sigma}} \lambda_n(\varepsilon^{-1}\xi) = \lambda_n^0(\xi) := |b_M(\xi)|^{\frac{1}{N}} e^{\frac{i\pi}{N}(\eta_M(\xi) + 2n)}, \text{ where } \sigma := \frac{N}{M}.$$

Proof. Recall from Remark 2.8 that the spectrum of the matrix $B(\varepsilon^{-1}\xi)$ consists of the solutions of the characteristic polynomial $\sum_{n=1}^N a_n \lambda^n = b(\varepsilon^{-1}\xi)$.

As $\{\alpha \in \mathbb{N}_M^d : |\alpha| = M \text{ and } b_\alpha \neq 0\} \neq \emptyset$, we see that $\varepsilon^M b(\varepsilon^{-1}\xi) \rightarrow b_M(\xi)$. Then, denoting by λ_ε an arbitrary solution of the last characteristic polynomial, we deduce that $(\lambda_\varepsilon)_{\varepsilon > 0}$ has no finite accumulation point. Together with the normalization $a_N = 1$, this in turn implies that $\sum_{n=1}^N a_n \lambda_\varepsilon^n \sim \lambda_\varepsilon^N$, and therefore

$$\lim_{\varepsilon \searrow 0} \varepsilon^M \lambda_\varepsilon^N = b_M(\xi) = |b_M(\xi)| e^{i\pi \eta_M(\xi)} := \lambda_N^0(\xi)^N,$$

which can be written equivalently as $\lim_{\varepsilon \searrow 0} \left(\varepsilon^{\frac{M}{N}} \frac{\lambda_\varepsilon}{\lambda_N^0(\xi)} \right)^N = 1$. Hence, the limiting spectrum of the matrix $B(\varepsilon^{-1}\xi)$ consists of N simple eigenvalues expressed in terms of the unit roots:

$$\lim_{\varepsilon \searrow 0} \varepsilon^{\frac{M}{N}} \lambda_n(\varepsilon^{-1}\xi) = \lambda_N^0(\xi) e^{\frac{2ni\pi}{N}} = \lambda_n^0(\xi), \quad n = 1, \dots, N.$$

□

The last result shows that the asymptotics of $B(\varepsilon^{-1}\xi)$ are in the context of Remark 2.8. Then, it follows that the space Fourier transforms of the Green functions are given by (2.12). Define the corresponding limits:

$$\hat{g}_n^0 := \Lambda_n^0 \sum_{j=1}^N \frac{e^{\lambda_j^0}}{\prod_{\ell \neq j} (\lambda_\ell^0 - \lambda_j^0)}, \quad n = 1, \dots, N, \quad (4.12)$$

with:

$$\Lambda_N^0(\xi) := 1, \quad \Lambda_n^0(\xi) := (-1)^{n-1} \sum_{\substack{1 \leq \ell_1 \leq \dots \leq \ell_{N-n} \leq N \\ \ell_1, \dots, \ell_{N-n} \neq n}} (\lambda_{\ell_1}^0 \cdots \lambda_{\ell_{N-n}}^0)(\xi) \quad \text{for } n < N.$$

The following additional condition is needed in order to characterize the short time asymptotics of the Green functions.

Assumption 4.6. *For all $\varphi \in \mathcal{S}$, the family $\{t^{1-n}\hat{g}_n(t, t^{-\sigma}\cdot)\varphi, t \in (0, \varepsilon]\}$ is uniformly integrable in $\mathbb{L}^1(\mathbb{R}^d)$, for some $\varepsilon > 0$.*

We denote by X the canonical map on \mathbb{R}^d , i.e. $X(x) = x$ for all $x \in \mathbb{R}^d$, and we introduce the scaled Green functions

$$g_n^\sigma(t, \cdot) := g_n(t, \cdot) \circ (t^\sigma X)^{-1}, \quad t \in [0, T], \quad n = 1, \dots, N.$$

In particular, in the case where g_n can be represented by a function, $g_n^\sigma(t, x) = t^{\sigma d} g_n(t, t^\sigma x)$.

Lemma 4.7. *The space Fourier Green functions \hat{g} satisfy the short time asymptotics*

$$t^{1-n}\hat{g}_n(t, t^{-\sigma}\xi) \longrightarrow \hat{g}_N^0(\xi) \quad \text{as } t \searrow 0, \text{ for all } \xi \in \mathbb{R}^d.$$

If in addition Assumption 4.6 holds true, then the last convergence holds in \mathcal{S}' , and the short time asymptotics of the scaled Green functions are given by

$$t^{1-n}g_n^\sigma(t, \cdot) \longrightarrow g_n^0 := \mathfrak{F}^{-1}\hat{g}_n^0 \quad \text{as } t \searrow 0, \text{ in } \mathcal{S}', \quad n = 1, \dots, N.$$

Proof. First, the pointwise convergence of $\hat{G}_n(t, \cdot) := t^{1-n}\hat{g}_n(t, t^{-\sigma}\cdot)$ towards \hat{g}_n^0 , as $t \searrow 0$, follows from direct application of Lemma 4.5 together with the observations reported in Remark 2.8. The uniform integrability condition of Assumption 4.6 guarantees that $\langle \hat{G}_n(t, \cdot), \varphi \rangle \longrightarrow \langle \hat{g}_n^0(t, \cdot), \varphi \rangle$ for all $\varphi \in \mathcal{S}$, i.e. $\hat{G}_n(t, \cdot) \longrightarrow \hat{g}_n^0$ as $t \searrow 0$ in \mathcal{S}' . This in turn implies the convergence of the corresponding Fourier inverse $\mathfrak{F}^{-1}\hat{G}_n(t, \cdot)$ towards $g_n^0 := \mathfrak{F}^{-1}\hat{g}_n^0$ as $t \searrow 0$ in \mathcal{S}' . It remains to relate the distribution $\mathfrak{F}^{-1}\hat{G}_n(t, \cdot)$ to the Green function g_n . To see this, we use the properties of the Fourier transform in \mathcal{S}' as defined by means of arbitrary test functions $\varphi \in \mathcal{S}$ as follows:

$$\langle \mathfrak{F}^{-1}\hat{G}_n, \varphi \rangle = \langle \hat{G}_n, \mathfrak{F}^{-1}\varphi \rangle = t^{\sigma d} \langle \hat{g}_n(t, \cdot), (\mathfrak{F}^{-1}\varphi)(t^\sigma \cdot) \rangle = t^{\sigma d} \langle g_n(t, \cdot), \mathfrak{F}((\mathfrak{F}^{-1}\varphi)(t^\sigma \cdot)) \rangle.$$

We finally observe by direct calculation that $\mathfrak{F}((\mathfrak{F}^{-1}\varphi)(\lambda\cdot)) = \lambda^{-d}\varphi(\lambda^{-1}\cdot)$ for all $\varphi \in \mathcal{S}$ and all constant $\lambda \in \mathbb{R}$. Then

$$\langle \mathfrak{F}^{-1}\hat{G}_n, \varphi \rangle = \langle g_n(t, \cdot), \varphi(t^{-\sigma}\cdot) \rangle = \langle g_n^\sigma(t, \cdot), \varphi \rangle.$$

By the arbitrariness of $\varphi \in \mathcal{S}$, this provides that $\mathfrak{F}^{-1}\hat{G}_n(t, \cdot) = g_n^\sigma(t, \cdot)$, thus completing the proof. \square

We observe that, as a consequence of the convergence of the scaled Green functions in \mathcal{S}' , we deduce that

$$t^{1-n+\sigma|\alpha|}\partial_x^\alpha g_n^\sigma(t, \cdot) \longrightarrow \partial_x^\alpha g_n^0 \quad \text{as } t \searrow 0, \text{ in } \mathcal{S}', \quad n = 1, \dots, N, \quad (4.13)$$

where $\partial_x^\alpha := \frac{\partial^{|\alpha|}}{\partial x_1^{\alpha_1} \dots \partial x_d^{\alpha_d}}$ for all $\alpha \in \mathbb{N}^d$. Our final result requires the following conditions.

- Assumption 4.8.** (i) *The functions $t \mapsto \int |z|^p |\gamma_n^1(t, z)|^p \mu_n^1(t, dz)$, $n = 1, \dots, N$, and $t \mapsto \int |\gamma_N^1(t, z)|^p \mu_N^1(t, dz)$ are continuous on $(0, T]$;*
(ii) $\gamma_n^1(t, t^\sigma z) = O(1)$ near the origin $t = 0$;
(iii) *The families*

$$\begin{aligned} & \left\{ z \mapsto |z|^p |\gamma_n^1(t, t^\sigma z)|^{p-1} |\partial_x g_n^\sigma(t, dz)| \right\}_{t \leq \varepsilon}, \quad 1 \leq n \leq N, \\ & \text{and } \left\{ z \mapsto |\gamma_N^1(t, t^\sigma z)|^{p-1} |\partial_x g_N^\sigma(t, dz)| \right\}_{t \leq \varepsilon} \end{aligned}$$

are uniformly integrable for some $\varepsilon > 0$.

Proposition 4.9. *Let Assumptions 2.1, 2.2, 4.1 (i), 4.6 and 4.8 hold true. Assume further that the density $\rho \in C^0(\mathbb{R}_+)$, strictly positive on $(0, \infty)$, and $\lim_{t \searrow 0} t^{N-1+\sigma p} \rho(t)^{1-p} < \infty$. Then, for any bounded Lipschitz function f , we have $\hat{r}_p + \hat{\alpha}_p < \infty$.*

Proof. By Assumption 4.1 (i), together with the fact that $\bar{\rho}(T) \leq \rho \leq 1$ on $[0, T]$, we only need to justify that

$$\sup_{\substack{0 \leq t \leq T \\ 1 \leq n \leq N}} \int |z|^p |\gamma_n^1(t, z)|^{p-1} |\partial_x g_n(t, dz)| + \sup_{0 \leq t \leq T} \rho(t)^{1-p} \int |\gamma_N^1(t, z)|^{p-1} |\partial_x g_N(t, dz)| < \infty.$$

Assumption 4.8 (i) ensures that the functions inside the last suprema are continuous on $(0, \infty)$. Then, in order to prove that \hat{r}_p and $\hat{\alpha}_p$ are finite, it suffices to verify that the functions inside the last suprema are bounded near $t = 0$.

To see this, we compute by a direct change of variables that:

$$\begin{aligned} \int |z|^p |\gamma_n^1(t, z)|^{p-1} |\partial_x g_n(t, dz)| &= \int |t^\sigma z|^p |\gamma_n^1(t, t^\sigma z)|^{p-1} |\partial_x g_n^\sigma(t, dz)| \\ &= t^{n-1+\sigma p} \int |z|^p |t^{1-n} \gamma_n^1(t, t^\sigma z)|^{p-1} |\partial_x g_n^\sigma(t, dz)| \\ &= O(t^{n-1+\sigma p}), \quad \text{near the origin,} \end{aligned}$$

by Assumption 4.8 (ii)-(iii), together with the consequence (4.13) of Lemma 4.7. This implies that the limit is zero as $n \geq 1$ and $p > 1$. Similarly,

$$\begin{aligned} \rho(t)^{1-p} \int |\gamma_N^1(t, z)|^{p-1} |\partial_x g_N|(t, dz) &= t^{N-1+\sigma p} \rho(t)^{1-p} \int t^{1-N} |\gamma_N^1(t, t^\sigma z)|^{p-1} |\partial_x g_N^\sigma|(t, dz) \\ &= O(t^{N-1+\sigma p} \rho(t)^{1-p}), \quad \text{near the origin,} \end{aligned}$$

again by Assumption 4.8 (ii)-(iii). \square

5 Numerical examples

5.1 Wave semi-linear PDE

We consider the nonlinear Klein-Gordon wave equation in \mathbb{R}^d for $1 \leq d \leq 3$:

$$(\partial_{tt}^2 - \Delta)u + u^3 + u^2 = 0, \quad u(0, x) = f_1(x), \quad \partial_t u(0, x) = f_2(x). \quad (5.1)$$

To the best of our knowledge, the current literature only considers approximation of the solution by deterministic numerical schemes, see e.g. [9]. Due to the curse of dimensionality, mainly $d = 1$ and $d = 2$ have been considered. To illustrate the efficiency of our algorithm, we solve this equation in $d = 1, 2$ and 3 . In our numerical experiments, we take the initial conditions

$$f_1(x) := -\frac{12}{9 + 2(\sum_{i=1}^d x_i)^2}, \quad \text{and} \quad f_2(x) := -\frac{48\sqrt{d+1}(\sum_{i=1}^d x_i)}{\left(2(\sum_{i=1}^d x_i)^2 + 9\right)^2},$$

for which the explicit solution is $u(t, x) = -\frac{12}{9 + 2(\sqrt{d+1}t - \sum_{i=1}^d x_i)^2}$.

We choose $\rho(t) = \beta e^{-\beta t}$ and $\bar{\rho}(t) = e^{-\beta t}$. For convenience, we set $u(t, x) := U(t, x) + f_1(x)$, and we compute that U satisfies the non-homogeneous nonlinear wave PDE:

$$\begin{aligned} (\partial_{tt}^2 - \Delta)U + U^3 + (3f_1 + 1)U^2 + (3f_1^2 + 2f_1)U + (f_1^3 + f_1^2 - \Delta f_1) &= 0 \\ U(0, x) = 0, \quad \partial_t U(0, x) &= f_2(x). \end{aligned}$$

Note that as $U(0, x) = 0$, we do not need to simulate our branching particles according to the distribution g_1 but only g_2 . This was our motivation for introducing the function U .

We directly compute that $r_p < (\frac{1+d}{6})^p t^p$ and $\alpha_p < t^p \beta^{1-p}$. Furthermore, $|(3f_1 + 1)|_\infty = 1$, $|(3f_1^2 + 2f_1)|_\infty = 8/3$ and $|(f_1^3 + f_1^2 - \Delta f_1)|_\infty < 1$. This implies that the assumptions in Theorem 3.5 are satisfied.

The Monte-Carlo approximation of $U(t, X_0)$ (and therefore $u(t, X_0) = U(t, X_0) + f_1(X_0)$) can be described by the following meta-algorithm:

Meta-algorithm:

1. Start at $t_0 = 0$ at the position X_0 and initialize a weight $\mathcal{W} := 1$.
2. Simulate an exponential r.v. τ with (arbitrarily) constant intensity β and simulate the particle at the new position $X_\tau = X_0 + Z_\tau$ at τ . More precisely, we draw uniform variables $(U_i)_{i=1,2}$ on $[0, 1]$ and set

$$\begin{aligned} Z_\tau^1 &= \tau(2U_1 - 1), \quad d = 1 \\ Z_\tau^1 &= \sqrt{1 - U_1^2} \cos(2\pi U_2)\tau, \quad Z_\tau^2 = \sqrt{1 - U_1^2} \sin(2\pi U_2)\tau, \quad d = 2 \\ Z_\tau^1 &= \cos(2\pi U_1) \cos(2\pi U_2)\tau, \quad Z_\tau^2 = \sin(2\pi U_1) \cos(2\pi U_2)\tau, \quad Z_\tau^3 = \sin(2\pi U_2)\tau, \quad d = 3. \end{aligned}$$

At $t_1 := \tau < t$, the particle dies and we create 0, 1, 2 or 3 descendants with probability $p := 1/4$. We then multiply the weight \mathcal{W} by the mass τ and according to the number of descendants, we update again the weight \mathcal{W} by

$$\mathcal{W} := \begin{cases} \mathcal{W} \times (-p^{-1})(f_1^3 + f_1^2 - \Delta f_1)(X_\tau)\beta^{-1}e^{\beta(t_1-t_0)}, & \text{if 0 descendant,} \\ \mathcal{W} \times (-p^{-1})(3f_1^2 + 2f_1)(X_\tau)\beta^{-1}e^{\beta(t_1-t_0)}, & \text{if 1 descendant} \\ \mathcal{W} \times (-p^{-1})(3f_1 + 1)(X_\tau)\beta^{-1}e^{\beta(t_1-t_0)}, & \text{if 2 descendants.} \\ \mathcal{W} \times (-p^{-1})\beta^{-1}e^{\beta(t_1-t_0)}, & \text{if 3 descendants,} \end{cases}$$

where 0 descendent means that the particle dies.

3. For each descendant, we apply independently Steps 2 and 3 until the default time – say τ_n – is greater than the maturity t . In this case, we multiply \mathcal{W} by

$$\mathcal{W} := \mathcal{W} \times e^{\beta(t-\tau_{n-1})}$$

4. Finally, for all particles alive at time t (with locations $(X_t^k)_{k \in \mathcal{K}_t}$), compute

$$\mathcal{W} \prod_{k \in \mathcal{K}_t} f_2(X_t^k),$$

and average the result using M Monte-Carlo paths.

In Figure 1, we have plotted two examples of Galton-Watson trees corresponding to the functionals:

$$\xi_{t, X_0}^{(1)} := \frac{1}{\beta^2 p^2} (3f_1(X_{\tau_1} + 1) e^{\beta\tau_1} e^{\beta(\tau_3-\tau_1)} e^{3\beta(t-\tau_3)} e^{\beta(t-\tau_1)} f_2(X_t^6) f_2(X_t^5) f_2(X_t^4) f_2(X_t^2)),$$

$$\xi_{t, X_0}^{(2)} := -\frac{1}{\beta^3 p^3} (3f_1(X_{\tau_4} + 1) (3f_1^2 + 2f_1)(X_t^6) e^{\beta\tau_1} e^{\beta(\tau_4-\tau_1)} e^{\beta(\tau_6-\tau_4)} f_2(X_t^7) f_2(X_t^5) f_2(X_t^3) f_2(X_t^2) e^{\beta(t-\tau_6)} e^{\beta(t-\tau_4)} e^{2\beta(t-\tau_1)}).$$

In the present very simple examples, in order to alleviate the figure, we have used simpler notations to label the branching particles than those in Section 3.1.

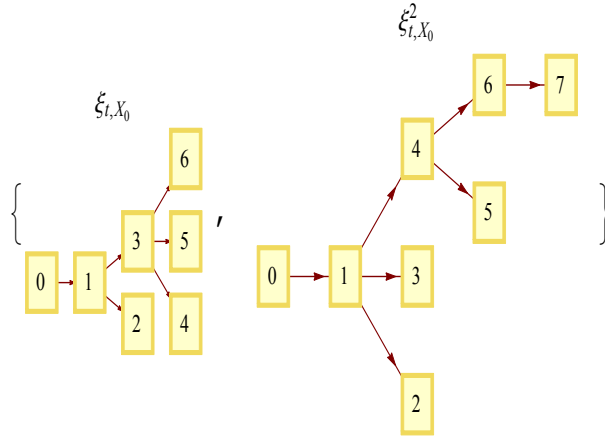


Figure 1: Galton-Watson trees associated to $\xi_{t,X_0}^{(1)}$ and $\xi_{t,X_0}^{(2)}$.

Note that by construction, the result is independent of β when M is large. There is an optimal choice of β that minimizes the variance of our Monte-Carlo estimator. We have chosen $\beta = 1$ in our numerical experiments and $M = 2^{22}$ for which the standard deviation of our estimator is less than 0.01. Below, we have plotted our numerical result for $u(t, X_0)$ against our analytical solution for $d = 1, 2, 3$ as a function of $x_0 \in [0, 1.5]$ (all the coordinates in \mathbb{R}^d are equal to x_0) and $t = 1$, see Figure 2. We obtain a perfect match. In order to see that our numerical solutions captures perfectly the additional nonlinearity $u^3 + u^2$, we have also shown for completeness the (analytical) solution of the linear wave equation (denoted “LinearKG”):

$$(\partial_{tt} - \Delta)u = 0, \quad u(0, x) = f_1(x), \quad \partial_t u(0, x) = f_2(x). \quad (5.2)$$

5.2 Yang-Mills PDE: a toy model

Here we consider the semilinear wave equation in \mathbb{R} :

$$\partial_{tt}u - \Delta u + u^3 + u\partial_x u = 0, \quad u(0, x) = f_1(x), \quad \partial_t u(0, x) = f_2(x). \quad (5.3)$$

We take the initial conditions

$$f_1(x) = -\frac{1}{1-x}, \quad f_2(x) = \frac{1}{(1-x)^2},$$

for which the explicit solution is $u(t, x) = -\frac{1}{1+t-x}$. Assumption (2.2) is satisfied for g_2 only for $d = 1$, this is why we restrict to this case. For $d > 1$, ∇g_2 involves derivative of delta function supported on the lightcone. Notice that the singularity of f_1 and f_2 is not

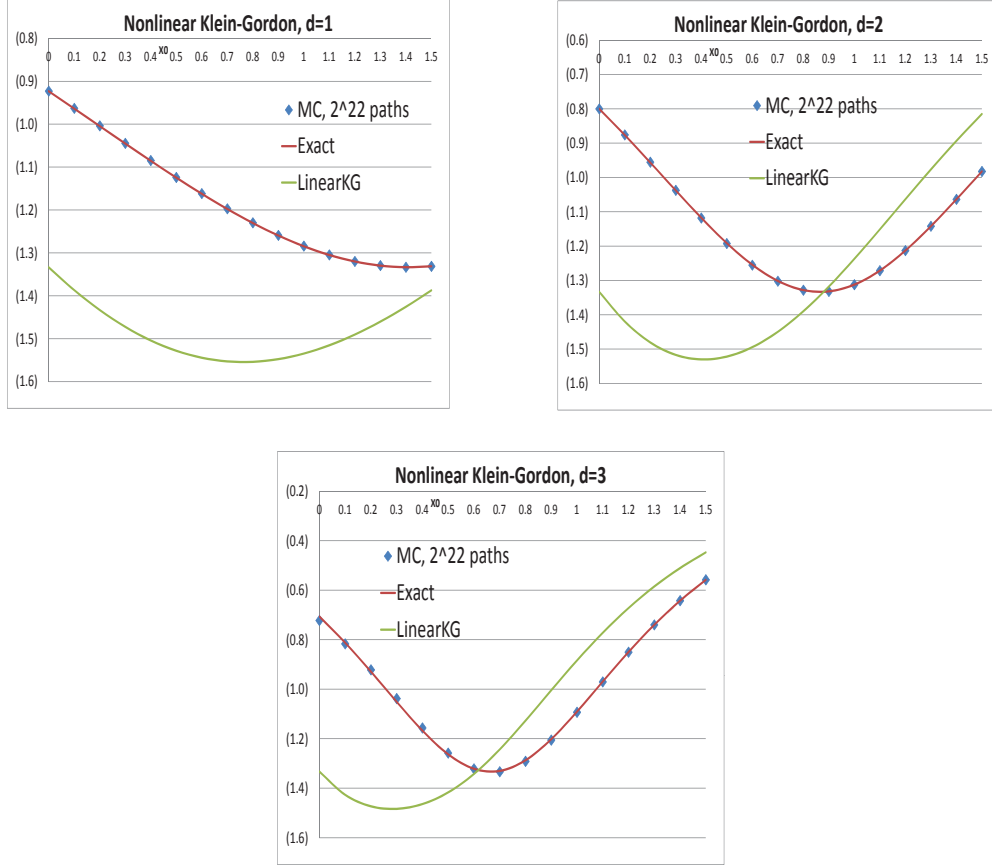


Figure 2: Numerical solutions of the nonlinear Klein-Gordon PDE (5.4) for $d = 1, 2, 3$ as a function of x_0 .

seen by our numerical algorithm because of our choice of the initial position x_0 and the finite speed property of the wave equation: The singularity $x = 1$ is not attainable.

This example can be interpreted as a scalar version of the Yang-Mills hyperbolic system. Indeed, the Lagrangian associated to the Yang-Mills theory on \mathbb{R}^{1+d} is

$$\int F_{\mu\nu}^a F^{\mu\nu a} d^4x, \quad F^{\mu\nu a} := \eta^{\mu\alpha} \eta^{\nu\beta} F_{\alpha\beta}^a$$

with $\eta_{00} = -1$, $\eta_{ii} = 1$, $i = 1, \dots, d$ and 0 otherwise. Here we use the convention of implicit summation for repeated indices (a, μ, ν) . The curvature is $F_{\mu\nu}^a := \partial_\mu A_\nu^a - \partial_\nu A_\mu^a - f_{bc}^a A_\mu^b A_\nu^c$. The Euler-Lagrange equations (written in the gauge frame $\partial_\mu A^{\mu,a} = 0$) give the system of hyperbolic PDE

$$\partial_\mu F^{\mu\nu a} + f_{bc}^a A_\mu^b F^{\mu\nu c} = 0$$

Note that the initial boundary conditions at $t = 0$, $A_\nu^a(t = 0, x)$ and $\partial_t A_\nu^a(t = 0, x)$ need to satisfy a constraint condition in order to preserve the gauge condition for all t . Using the expression for the curvature $F_{\mu\nu}^a$, we get that $A^{\mu a} := \eta^{\mu\alpha} A_\alpha^a$ is solution of a system of hyperbolic PDE of the form:

$$\square A^{\nu a} - f_{bc}^a A^{\mu b} \partial_\mu A^{\nu c} + f_{bc}^a A_\mu^b \left(\partial^\mu A^{\nu c} - \partial^\nu A^{\mu c} - f_{de}^c A^{\mu d} A^{\nu e} \right) = 0$$

with $\square := \partial_{tt} - \Delta$. This can schematically written as

$$\square A - f A \partial A - f^2 A^3 = 0$$

hence our PDE (5.3).

Note that our branching Monte-Carlo algorithm can be easily adapted to solve a system of semilinear PDEs. All we need to do is to index particles with a type (μa) corresponding to a coordinate of the solution, see an example with the complex Gross-Pitaevskii equation in Section 5.4.

As in the previous section, we set $u(t, x) = f_1(x) + U(t, x)$ for which U satisfies the non-homogeneous hyperbolic PDE

$$\begin{aligned} \partial_{tt} U - \Delta U + U^3 + 3f_1 U^2 + U \sum_{i=1}^d \partial_{x_i} U + (3f_1^2(x) + \sum_{i=1}^d \partial_{x_i} f_1) U + f_1 \sum_{i=1}^d \partial_{x_i} U \\ + (f_1^3 + \sum_{i=1}^d f_1 \partial_{x_i} f_1 - \Delta f_1) = 0, \quad U(0, x) = 0, \quad \partial_t u(0, x) = f_2(x) \end{aligned}$$

The Monte-Carlo approximation of $U(t, X_0)$ (and therefore $u(t, X_0) = U(t, X_0) + f_1(X_0)$) can be described by the following meta-algorithm:

1. Start with a type (0) particle at $t_0 = 0$ at position X_0 , and initialize a weight $\mathcal{W} := 1$;
2. Simulate an exponential r.v. τ with (arbitrarily) constant intensity β , and simulate the particle at the new position $X_\tau = X_0 + Z_\tau$ at τ . More precisely, we draw a uniform variable (U) on $[0, 1]$ and set $Z_\tau = \tau(2U - 1)$;
3. At $t_1 := \tau < t$, the particle dies and we create 0, 1, 2 or 3 descendants with probability $p := 1/6$; in case of 1 descendant, the type assigned is (0) or (1); similarly, in case of 2 descendants, their types can be both (0), both (1), or one of each type; in case of 3 descendants, they are all of type (0); we then multiply the weight \mathcal{W} by the mass τ (see Example 2.10-(i)) if the type of the particle is (0) and according to the number of descendants and type, we update also the weight \mathcal{W} by

$$\mathcal{W} := \begin{cases} \mathcal{W} - p^{-1}(f_1^3 + f_1 \partial_x f_1 - \Delta f_1)(X_\tau) \beta^{-1} e^{\beta(t_1 - t_0)}, & 0 \text{ descendant,} \\ \mathcal{W} \times -p^{-1}(3f_1^2(x) + \partial_x f_1)(X_\tau) \beta^{-1} e^{\beta(t_1 - t_0)}, & 1 \text{ descendant, type (0)} \\ \mathcal{W} \times -p^{-1} f_1(X_\tau) \beta^{-1} e^{\beta(t_1 - t_0)}, & 1 \text{ descendant, type (1)} \\ \mathcal{W} \times (-3/p) f_1(X_\tau) \beta^{-1} e^{\beta(t_1 - t_0)}, & 2 \text{ descendants, type (0)} \\ \mathcal{W} \times -p^{-1} \beta^{-1} e^{\beta(t_1 - t_0)}, & 2 \text{ descendants, type (0) and (1)} \\ \mathcal{W} \times -p^{-1} \beta^{-1} e^{\beta(t_1 - t_0)}, & 3 \text{ descendants,} \end{cases}$$

where 0 descendant means that the particle dies. The particle of type (1) is then simulated on the light cone, meaning that $Z_\tau = \tau$ or $Z_\tau = -\tau$ with probability $1/2$. In the last case, the weight is multiplied by -1 .

4. For each particle, we apply independently Steps 2 and 3 until the default time – say τ_n – is greater than the maturity t . In this case, we multiply \mathcal{W} by

$$\mathcal{W} := \mathcal{W} \times e^{\beta(t-\tau_{n-1})}$$

5. Finally, compute for all particles in \mathcal{K}_t

$$\mathcal{W} \prod_{k \in \mathcal{K}_t} \left(f_2(X_t^k) 1_{\text{type}=0} - f_2(X_t^k - \Delta X_t^k) 1_{\text{type}=1} \right)$$

where $\Delta X_t^k = \pm(t - T_{n-1})$ with probability $1/2$ and average the result using M Monte-Carlo paths.

Below, we have plotted our numerical result for $u(t, X_0)$ again our analytical solution for $d = 1$ as a function of $x_0 \in [3, 5]$ and $t = 1$, see Figure 4. We obtain a perfect match. For completeness, we have also shown the (analytical) solution of the linear wave equation (denoted “Linearwave”)

$$(\partial_{tt} - \Delta)u = 0, \quad u(0, x) = f_1(x), \quad \partial_t u(0, x) = f_2(x) \tag{5.4}$$

to show that our numerical solutions capture perfectly the additional nonlinearity $u^3 + u\partial_x u$. Note that as $X_t \in [x_0 - t, x_0 + t]$, the singularity of f_2 at $x = 1$ is not relevant.

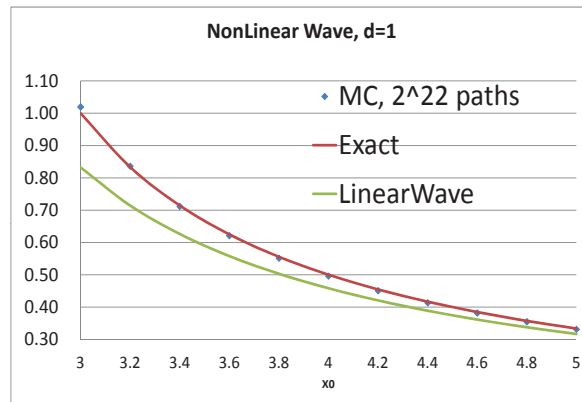


Figure 3: Numerical solution of the 1-d semi-linear wave PDE (5.3) as a function of x_0 .

5.3 Nonlinear Beam PDE

We consider the nonlinear beam equation:

$$\partial_t^2 u + \partial_x^4 u + u^2 + h(t, x) = 0, \quad u(0, x) = \tanh(x), \quad \partial_t u(0, x) = \cosh(x)^{-2}, \quad (5.5)$$

for which the explicit solution is $u(t, x) = \tanh(x + t)$ for a suitable choice of h . Here we follow the same discussion as in Section 5.1. Below, we have plotted our numerical result for $u(t, X_0)$ against our exact solution for $d = 1$ as a function of $x_0 \in [-0.5, 0.5]$ and $t = 0.5$, see Figure 4. Here we obtain a small error due to the fact that the one-dimensional density $\frac{|G(z)|}{\int_{\mathbb{R}} |G(z)| dz}$ with G given by (2.11) has been splined on an interval $[-10, 10]$ for computational purpose.

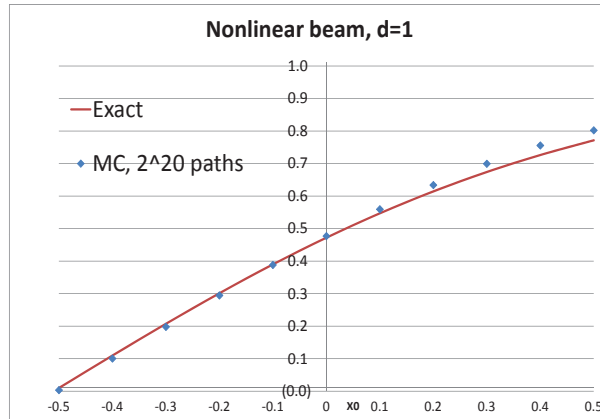


Figure 4: Numerical solution of nonlinear beam PDE (5.5) for $d = 1$ as a function of x_0 .

5.4 Gross-Pitaevskii PDE

The Gross-Pitaevskii PDE reads

$$i\partial_t u(t, x) = -\frac{1}{2}\Delta u(t, x) + h|u(t, x)|^2 u(t, x), \quad x \in \mathbb{R}^d \quad (5.6)$$

with h a constant. This equation describes a Bose-Einstein condensate at zero or very low temperature. This has been recently solved using a time-splitting spectral method [4]. This deterministic method suffers from the curse of dimensionality and requires suitable mesh size controls. Below, we present our Monte-Carlo algorithm. We should again emphasize that our algorithm is immune to the dimension and if the standard deviation of our Monte-Carlo estimate converges to zero, we can ensure that we have converged to the true solution.

We set $h := -1$ and $f_1(x) := \frac{\sqrt{d}}{\cosh(\sum_{i=1}^d x_i)}$ for which the explicit solution is

$$u(t, x) = e^{\frac{idt}{2}} \frac{\sqrt{d}}{\cosh(\sum_{i=1}^d x_i)}$$

PDE (5.6) can be written as a two-dimensional PDE system with polynomial nonlinearity:

$$\begin{aligned} i\partial_t u(t, x) &= -\frac{1}{2}\Delta u + guu^*u \\ -i\partial_t u^*(t, x) &= -\frac{1}{2}\Delta u^* + guu^*u^* \end{aligned}$$

From Example 2.11, $Z_\tau = e^{\frac{i\pi}{4}}\sqrt{\tau}Z \in \mathbb{C}^d$ and $Z_\tau^* = e^{-\frac{i\pi}{4}}\sqrt{\tau}Z \in \mathbb{C}^d$ with $Z \in N(0, d)$.

Meta-algorithm:

1. Start at $t_0 = 0$ at the position X_0 with a particle of type 0 and initialize a *complex* weight $\mathcal{W} := 1$.
2. Simulate an exponential r.v. τ with (arbitrarily) constant intensity β and simulate the particle at the new (complex) position $X_\tau = X_0 + e^{\frac{i\pi}{4}}\sqrt{\tau}Z$ (resp. $X_\tau = X_0 + e^{-\frac{i\pi}{4}}\sqrt{\tau}Z$) at τ if the particle is of type 0 (resp. of type 1) with Z a d -dimensional standard (real) Gaussian variable.

At $t_1 := \tau < t$, the particle dies and we create 2 descendants of type 0 and one of type 1 (resp. 2 descendants of type 1 and one of type 0) if the particle is of type 0 (resp. type 1). We then multiply the weight \mathcal{W} by

$$\mathcal{W} := \begin{cases} \mathcal{W} \times (i)\beta^{-1}e^{\beta(t_1-t_0)}, & \text{type 0} \\ \mathcal{W} \times (-i)\beta^{-1}e^{\beta(t_1-t_0)}, & \text{type 1} \end{cases}$$

3. For each descendant, we apply independently Steps 2 and 3 until the default time – say τ_n – is greater than the maturity t . In this case, we multiply \mathcal{W} by

$$\mathcal{W} := \mathcal{W} \times e^{\beta(t-\tau_{n-1})}$$

4. Finally, for all particles alive at time t (with locations $(X_t^k)_{k \in \mathcal{K}_t}$), compute

$$\mathcal{W} \prod_{k \in \mathcal{K}_t : \text{Type}=0} f_1(X_t^k) \prod_{k \in \mathcal{K}_t : \text{Type}=1} f_1(X_t^k)^*,$$

and average the result using M Monte-Carlo paths.

We have chosen $\beta = 1$ in our numerical experiments and $M = 2^{22}$ for which the standard deviation of our estimator is less than 0.01. Below, we have plotted our numerical result for $Re(u(t, x_0))$ and $Im(u(t, x_0))$ against our analytical solution for $d = 1, 2, 3$ as a function of $x_0 \in [0, 1.5]$ (all the coordinates in \mathbb{R}^d are equal to x_0) and $t = 0.1$, see Figures 5, ???. We obtain a perfect match. For completeness, we have also shown the (numerical) solution of the linear Schrodinger equation (i.e., $g := 0$), denoted “Linear”.

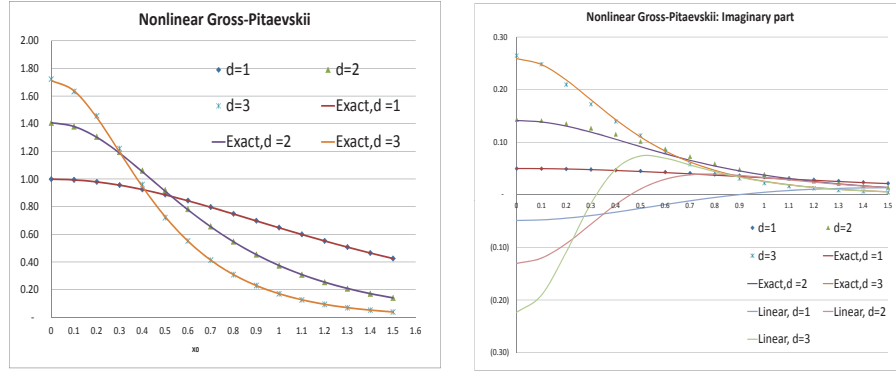


Figure 5: Numerical solutions for $Re(u(t, x_0))$ (left) and $Im(u(t, x_0))$ (right) of the Gross-Pitaevskii PDE (5.6) for $d = 1, 2, 3$ as a function of x_0 .

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